# SPECTRAL ANALYSIS OF DISCRETE-TIME SIGNALS

# Random discrete signals

- Exact values are unknown outside of the range in which the signal is observed
- Described through statistical parameters
  - This approach is sensible for many real-world signals, which occur as a consequence of mutual interaction of a number a complex causes
- A random signal is one specific realization of a *random process* 
  - A discrete-time random process X<sub>n</sub> represents a time series of random variables

# Statistical ensemble

- A set of all possible realizations of a random process
  - different signals but same statistical properties



## Random variables and random processes

 A random variable can be described through its *cumulative probability distribution function* (CDF):

 $F_X(x) = P(X \leq x)$ 

and if it is continuous, it can also be described through its *probability density function* (PDF):

$$p_{X}(x) = \frac{dF_{X}(x)}{dx}$$

 In this way we can also describe any random process at a particular moment



## Random processes

A random process at a moment n can be described using a CDF dependent on n:

$$F_{X_n}(x,n) = \mathsf{P}(X_n \leq x)$$

and if it is continuous, it can also be described using a PDF dependent on n:

$$p_{X_n}(x,n) = \frac{dF_{X_n}(x,n)}{dx}$$

• The values of X<sub>n</sub> are generally mutually dependent, and to fully describe a random process we should know a joint CDF for any subset of moments:

$$F_{X_{n_1}X_{n_2}...X_{n_N}}(x_1, x_2, ..., x_N, n_1, n_2, ..., n_N) = P(X_{n_1} \le x_1, X_{n_2} \le x_2, ..., X_{n_N} \le x_N)$$

 In practice this function is most often unknown, and we describe random processes using statistical parameters such as mean and autocorrelation function

## Stationarity of a random process

• A random process is (*strict-sense*) *stationary* if its joint CDF does not change in time, i.e. if, for any integer *k*, the following holds:

$$F_{X_{n_1}X_{n_2}...X_{n_N}}(x_1, x_2, ..., x_N, n_1, n_2, ..., n_N)$$
  
=  $F_{X_{n_1+k}X_{n_2+k}...X_{n_N+k}}(x_1, x_2, ..., x_N, n_1 + k, n_2 + k, ..., n_N + k)$ 

 There are many random processes for which this does not hold, and which go through various *states*, in which their statistical properties may be quite different

### Mean and autocorrelation function

• The *mean* of a random process is its mathematical expectation:

$$m_{X_n}(n) = E\{X_n\} = \int_{-\infty}^{\infty} x p_{X_n}(x, n) dx$$

and if it does not depend on *n*, the process is *stationary with respect to its mean* 

• The *autocorrelation function* of a random process is the measure of the similarity of its values at the time instants *n* and *n*+*k*:

$$r_{X_n X_{n+k}}(n, n+k) = E\{X_n X_{n+k}\} = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} x_1 x_2 p_{X_n X_{n+k}}(x_1, x_2, n, n+k) dx_1 dx_2$$

and if it depends solely on k, the process is stationary with respect to its autocorrelation function

• A process stationary both in terms of its mean and its autocorrelation function is considered *wide-sense stationary* 

### Variance

• The variance of a wide-sense stationary (WSS) random process is the measure of its expected deviation from its mean:

If m<sub>x</sub> = 0, the variance is equal to the *average power* of the signal

### Estimation of statistical parameters

- We usually have only one member of the statistical ensemble x(n) at our disposal, from which we should infer the values of the statistical parameters of a WSS random process
- The mean of the discrete signal x(n) is:

$$\overline{x(n)} = \lim_{N \to \infty} \frac{1}{2N+1} \sum_{n=-N}^{N} x(n)$$

and its autocorrelation is:

$$r_{xx}(k) = \lim_{N \to \infty} \frac{1}{2N+1} \sum_{n=-N}^{N} x(n) x(n+k)$$

and if the process is *ergodic*, they are equal to the ensemble mean and autocorrelation function

 In practice we cannot calculate either of these parameters because we would need to know infinitely many values of signal samples and calculate infinite sums

### Examples of parameter estimation

#### Mean estimation

• The standard estimate of the mean is:

$$\hat{m}_x = \frac{1}{N} \sum_{n=0}^{N-1} x(n)$$

This estimate is *unbiased* (its expectation is equal to the actual mean m<sub>x</sub>):

$$E\{\hat{m}_{x}\} = \frac{1}{N}E\left\{\sum_{n=0}^{N-1}X_{n}\right\} = \frac{1}{N}\sum_{n=0}^{N-1}E\{X_{n}\} = m_{X}$$

- This estimate is also *consistent* (variance goes to 0 when N goes to infinity)
  - Consistency is a highly desirable property of an estimate because with a consistent estimate we can expect the same result from any member of the ensemble

### Examples of parameter estimation

#### **Estimation of the autocorrelation function**

• The standard estimate of the autocorrelation function is:

$$\hat{r}_{xx}(k) = \frac{1}{N} \sum_{n=0}^{N-|k|-1} x(n) x(n+k)$$

This estimate is *biased* (its expectation is not equal to the actual autocorrelation function r<sub>xx</sub>(k)):

$$E\{\hat{r}_{xx}(k)\} = \frac{1}{N} \sum_{n=0}^{N-|k|-1} E\{X_n X_{n+k}\} = \frac{1}{N} \sum_{n=0}^{N-|k|-1} r_{xx}(k) = \left(1 - \frac{|k|}{N}\right) r_{xx}(k)$$

but it is asymptotically unbiased (its expectation converges to  $r_{xx}(k)$  as N approaches infinity)

• This estimate is *consistent* 

### Spectrum of a discrete-time random process

- Discrete-time random signals have infinite energy and thus have no DTFT, and we cannot calculate the energy spectral density (which is equal to the square of the magnitude spectrum)
- Discrete-time random signals have finite *average power*:

$$P_{x} = \lim_{N \to \infty} \frac{1}{2N+1} \sum_{n=-N}^{N} |x(n)|^{2}$$

and it is possible to define their *power spectral density* (*power spectrum*), by analogy with a corresponding definition for deterministic signals:

$$S_{X}(\omega) = \lim_{N \to \infty} E\left\{\frac{1}{2N+1}\left|\sum_{n=-N}^{N} X_{n} e^{-j\omega n}\right|^{2}\right\}$$

### Spectrum of a discrete-time random process

• Power spectral density (PSD) is equal to the Fourier transform of the autocorrelation function:

$$S_{\chi}(\omega) = \sum_{k=-\infty}^{\infty} r_{\chi\chi}(k) e^{-j\omega k}$$

and it is a real, non-negative function, and in the case of real signals also even

- The estimation of power spectral density is one of the most frequent and most important tasks in digital signal processing
  - The problem is that we usually have only one member of the statistical ensemble at our disposal, and we usually know the values of only a limited number of its samples

## Methods for power spectrum estimation

#### Non-parametric (classical) methods

- The spectrum is directly calculated by applying DTFT to either a finite-length segment of the input singal or its autocorrelation
- We do not introduce any assumptions as to the shape of the power spectrum

#### **Parametric methods**

- We introduce a *model* of power spectral density function and the problem amounts to the estimation of the parameters of this model
- If an adequate model is used, these methods can give much better results than non-parametric ones, particularly if only a short segment of the input signal is available

### Non-parametric methods for estimating $S_{\chi}(\omega)$

• The estimation of  $S_{\chi}(\omega)$  is more accurate if a longer segment of the input signal is available (as was the case with deterministic signals)



- If *N* is too small, we can increase the density of DFT samples by zero padding, although this does not affect frequency resolution
  - Zero padding, as a means for increasing the density of DFT samples is of particular importance for non-stationary signals

## Periodogram

• Direct estimate based on the definition of  $S_{\chi}(\omega)$ , with the time range limited to the interval from 0 to N-1:

$$S_{x}(\omega) = \lim_{N \to \infty} E\left\{\frac{1}{2N+1}\left|\sum_{n=-N}^{N} X_{n} e^{-j\omega n}\right|^{2}\right\} \rightarrow \left\|\hat{S}_{x}(\omega) = \frac{1}{N}\left|\sum_{n=0}^{N-1} x(n) e^{-j\omega n}\right|^{2}\right\}$$

which is a real, non-negative function, and in the case of real signals also even, just as was the case for  $S_x(\omega)$ 

### Periodogram

$$\hat{S}_{x}(\omega) = \frac{1}{N} |\bar{X}(\omega)|^{2} = \frac{1}{N} \bar{X}(\omega) \bar{X}^{*}(\omega)$$

$$= \frac{1}{N} \sum_{n=0}^{N-1} x(n) e^{-j\omega n} \sum_{m=0}^{N-1} x(m) e^{j\omega m}$$

$$= \frac{1}{N} \sum_{n=0}^{N-1} \sum_{m=0}^{N-1} x(n) x(m) e^{-j\omega(m-n)}$$

$$= \sum_{k=-N+1}^{N-1} \left( \frac{1}{N} \sum_{n=0}^{N-|k|-1} x(n) x(n+k) \right) e^{-j\omega k}$$

$$= \sum_{k=-N+1}^{N-1} \hat{r}_{xx}(k) e^{-j\omega k}$$

• Periodogram represents the Fourier transform of the standard estimate of the autocorrelation function of a random process

### Properties of the periodogram

$$E\left\{\hat{S}_{x}(\omega)\right\} = E\left\{\sum_{k=-N+1}^{N-1}\hat{r}_{xx}(k)e^{-j\omega k}\right\} = \sum_{k=-N+1}^{N-1}E\left\{\hat{r}_{xx}(k)\right\}e^{-j\omega k}$$
$$= \sum_{k=-N+1}^{N-1}\left(1 - \frac{|k|}{N}\right)r_{xx}(k)e^{-j\omega k}$$

- The expectation of the periodogram  $\hat{S}_{\chi}(\omega)$  is not equal to the actual power spectrum  $S_{\chi}(\omega) = \text{DTFT}\{r_{\chi\chi}(k)\}$ , but to  $\text{DTFT}\{r_{\chi\chi}(k)w_B(k)\}$ , where  $w_B(k)$  is the triangular window function
  - Periodogram is a *biased* estimate of  $S_{\chi}(\omega)$ , but still *asymptotically unbiased*
- When estimating  $S_{\chi}(\omega)$  on the basis of a periodogram, the problem of *spectral leakage* occurs

$$E\left\{\hat{S}_{x}(\omega)\right\} = S_{x}(\omega) \circledast W_{B}(\omega) \qquad W_{B}(\omega) = \frac{1}{N} \left(\frac{\sin N\omega/2}{\sin \omega/2}\right)^{2}$$

### Properties of the periodogram

• Variance of the periodogram  $\hat{S}_{\chi}(\omega)$  for large values of N in most cases of practical interest is proportional to the square of  $S_{\chi}(\omega)$ 

$$\lim_{N\to\infty} \operatorname{Var}\left\{\hat{S}_{x}(\omega)\right\} \sim S_{x}^{2}(\omega)$$

- Periodogram is an *inconsistent* estimate of  $S_X(\omega)$ , which means that it can greatly depend on the realization of the random process
- Due to its inconsistency, periodogram is not considered to be a high quality estimate of  $S_x(\omega)$
- An idea: to split the available segment of the input signal into K equal parts, to calculate the periodogram on each of them and to calculate the average of the results

## Periodogram averaging

- Let the available segment of the input signal x(n) of length N be divided into K equal parts of length M = N/K, with no overlap (basic version, proposed by Bartlett)
- Let  $\hat{S}_{x}^{(i)}(\omega)$  be the periodogram calculated on the *i*-th segment:

$$\hat{S}_{x}^{(i)}(\omega) = \frac{1}{M} \left| \sum_{n=0}^{M-1} x_{i}(n) e^{-j\omega n} \right|^{2}$$

• The averaged periodogram is obtained as the average of thus obtained *K* periodograms:

$$\hat{S}_x^B(\omega) = \frac{1}{\kappa} \sum_{i=0}^{\kappa-1} \hat{S}_x^{(i)}(\omega)$$

## Properties of the averaged periodogram

• Mathematical expectation of  $\hat{S}_{x}^{B}(\omega)$  is equal to the mathematical expectation of a single periodogram  $\hat{S}_{x}^{(i)}(\omega)$ :

$$E\left\{\hat{S}_{x}^{B}(\omega)\right\} = \sum_{k=-M+1}^{M-1} \left(1 - \frac{|k|}{M}\right) r_{XX}(k) e^{-j\omega k} = S_{X}(\omega) \circledast W_{B}(\omega)$$

and since the triangular window function is now *K* times shorter, the frequency resolution is also *K* times worse

• The variance is reduced by a factor of *K* with respect to the standard periodogram (the variance of the mean of *K* identically distributed random variables is *K* times smaller than the variance of a single one of them):

$$\lim_{v\to\infty} \operatorname{Var}\left\{\hat{S}_{x}(\omega)\right\} \sim \frac{1}{K}S_{x}^{2}(\omega)$$

• By averaging in this way, we have improved the consistency of the estimate at the expense of the frequency resolution

# Modifications of the averaged periodogram

- The basic idea of Bartlett can be modified in two ways (both proposed by Welch):
  - Consecutive segments can overlap
    - This increases the number of segments and decreases variance further
    - Overlap should not be too significant because that would undermine the assumption of independence between consecutive segments (if the overlap is above 50% the variance will abruptly stop decreasing)
  - Every segment x<sub>i</sub>(n) can be shaped by applying a non-rectangular window function before calculating the periodogram
    - This reduces the correlation between consecutive segments, and also mitigates the problem of lower reliability of the estimate of autocorrelation function for large k
- Without affecting the frequency resolution, the variance is now lower: for a large N it still behaves as  $S_x^2(\omega)/K$ , but K is now greater owing to the overlap

### Periodogram smoothing

- The idea is based on windowing the estimate of the autocorrelation function (as proposed by Blackman and Tukey):
  - This reduces the influence of less reliable values of  $\hat{r}_{xx}(k)$ , obtained for large |k| $\hat{c}^{BT}(k) = \sum_{k=1}^{K-1} \hat{c}_{k}(k) = \frac{k^{-1}}{2} \hat{c$

$$\hat{S}_{x}^{BT}(\omega) = \operatorname{FTD}\left\{\hat{r}_{xx}(k)w(k)\right\} = \sum_{k=-K+1}^{K-1}\hat{r}_{xx}(k)w(k)e^{-j\omega k}$$

• Mathematical expectation of a thus smoothed periodogram is:

$$E\left\{\hat{S}_{x}^{BT}(\omega)\right\} = \operatorname{FTD}\left\{E\left\{\hat{r}_{xx}(k)\right\}w(k)\right\} = \operatorname{FTD}\left\{\hat{r}_{xx}(k)w_{B}(k)w(k)\right\}$$
$$= S_{x}(\omega) \circledast W_{B}(\omega) \circledast W(\omega) = E\left\{\hat{S}_{x}(\omega)\right\} \circledast W(\omega)$$

where w(k) is usually much shorter than  $w_B(k)$ , and thus the influence of convolution with  $W_B(\omega)$  is negligible

- The resulting spectral leakage depends on  $W(\omega)$ 
  - The choice of W(ω) affects the trade-off between the variance and the frequency resolution

# Parametric methods for estimating $S_X(\omega)$

- We introduce a *model* of the random process, and the problem of estimating  $S_{\chi}(\omega)$  amounts to the problem of estimating model parameters
- We assume that the signal x(n) is the output of an LTI system with transfer function:

$$H(z) = \frac{B(z)}{A(z)} = \frac{\sum_{i=0}^{q} b_i z^{-i}}{1 + \sum_{i=1}^{p} a_i z^{-i}}$$

where the input is white noise s(n) of zero mean and variance (average power) equal to  $\sigma^2$ 

The term white implies that the samples of noise are mutually uncorrelated, i.e.:

$$r_{ss}(k) = \sigma^2 \delta(k)$$

- This reduces the problem to the modeling of a system
- Power spectral density of the output signal is equal to:  $S_{\chi}(\omega) = \sigma^2 \frac{|B(\omega)|^2}{|\Delta(\omega)|^2}$

# Parametric methods for estimating $S_{\chi}(\omega)$

- First we need to choose the values for *p* and *q*, and then to estimate *a<sub>i</sub>* and *b<sub>i</sub>* 
  - q = 0: AR (autoregressive) models
  - *p* = 0: MA (*moving average*) models
  - $q \neq 0, p \neq 0$ : ARMA models



- The choice is usually based on some previous knowledge of the random process
  - Owing to their simplicity, AR models are most commonly used (the others yield non-linear systems of equations)
  - An example of a problem where AR is the obvious solution is the signal which is composed of a sinusoid of an unknown frequency and noise
    - Poles of the model should match the unknown frequency of the sinusoid

### Estimation of $S_{\chi}(\omega)$ for AR random processes

$$\begin{aligned} x(n) &= -\sum_{i=1}^{p} a_{i} x(n-i) + s(n) \\ & E\{x(n)x(n-k)\} = -\sum_{i=1}^{p} a_{i} E\{x(n-i)x(n-k)\} + E\{s(n)x(n-k)\} \\ & r_{xx}(k) = -\sum_{i=1}^{p} a_{i} r_{xx}(k-i) + E\{s(n)\sum_{m=0}^{\infty} h(m)s(n-k-m)\} \\ & = -\sum_{i=1}^{p} a_{i} r_{xx}(k-i) + \sum_{m=0}^{\infty} h(m) E\{s(n)s(n-k-m)\} \\ & = -\sum_{i=1}^{p} a_{i} r_{xx}(k-i) + \sum_{m=0}^{\infty} h(m)\sigma^{2}\delta(m+k) & h(m)\sigma_{x}(k-i) + \sum_{m=0}^{\infty} h(m)\sigma^{2}\delta(m+k) \\ & = \begin{cases} -\sum_{i=1}^{p} a_{i} r_{xx}(k-i) + \sigma^{2}, & k = 0 \end{cases} & \text{we hat follow} \end{aligned}$$

we assume  $k \ge 0$ , and the values for negative k can be easily found because  $r_{xx}(k)$  is even

h(m) exists only for  $m \ge 0$ , and thus for  $k \ge 0$  the only case when the right sum is non-zero is when k = m = 0

we have used h(0) = 1, which follows from B(z) = 1

### Estimation of $S_{\chi}(\omega)$ for AR random processes

• Coefficients  $a_i$  and  $\sigma^2$  are the solutions of the Yule-Walker equations:

$$\begin{bmatrix} r_{XX}(0) & r_{XX}(-1) & \dots & r_{XX}(-p) \\ r_{XX}(1) & r_{XX}(0) & \dots & r_{XX}(-p+1) \\ \vdots & \vdots & \ddots & \vdots \\ r_{XX}(p) & r_{XX}(p-1) & \dots & r_{XX}(0) \end{bmatrix} \begin{bmatrix} 1 \\ a_1 \\ \vdots \\ a_p \end{bmatrix} = -\begin{bmatrix} \sigma^2 \\ 0 \\ \vdots \\ 0 \end{bmatrix}$$

- The system matrix is a symmetrical Toeplitz matrix (each descending diagonal from left to right is constant), and thus the system can be solved efficiently, in N<sup>2</sup> instead of N<sup>3</sup> time, using the Levinson-Durbin algorithm
  - The same matrix is obtained if the problem is posed as a *linear prediction* problem, i.e. if the following is assumed:

$$\hat{x}(n) = -\sum_{i=1}^{p} a_i x(n-i)$$

 In practice the autocorrelation function r<sub>xx</sub>(k) is not available and we only have its estimate, but the values of this estimate for k ≤ p are still much more reliable than the values for relatively large k